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GRM: Observing
the Terrestrial Gravity
and Magnetic Fields
in the 1990's

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Beginning with the earliest days of space exploration, satellites have been used to chart the gravity and magnetic fields of the earth. As a continuation of these studies NASA is proposing a launch a new geopotential fields exploration system called the Geopotential Research Mission (GRM). Two spacecraft will be placed in a circular polar orbit at 100 km altitude. Distances between these satellites will vary from 100 to 600 km. Both scalar and vector magnetic fields will be measured by magnetometers mounted on a boom positioned in the forward direction on the lead satellite. Gravity data will be computed from the measured change in distance between the two spacecraft. This quantity, called the range-rate, will be determined from the varying frequency (Doppler shift) between transmitter and receiver on each satellite. Expected accuracies (at the one sigma level) are: gravity field, $1 \times 10^{-5} \text{ m s}^{-2}$ (1 mGal); 5 cm geoid height; magnetic, scalar field 5 nT, vector to 20 arc seconds (106 microrad), both resolved to less than 100 km.

With these more accurate and higher resolution data we will be able to investigate the earth's structure from the crust (with the shorter wavelength gravity and magnetic anomalies) through the mantle (from the intermediate wavelength gravity field) and into the core (using the longer wavelength gravity and magnetic fields).

Introduction

From the very beginning of the artificial satellite era, space platforms have been used to map the gravity and magnetic fields of the earth. In 1958 Vanguard 1 (O'Keefe et al., 1958) and Sputnik 3 were the first satellites to map these geopotential fields. Tables 1 and 2 list the most significant satellites that have been used to study the near-earth gravity and magnetic fields. The major difference between these two groups of satellites (Tables

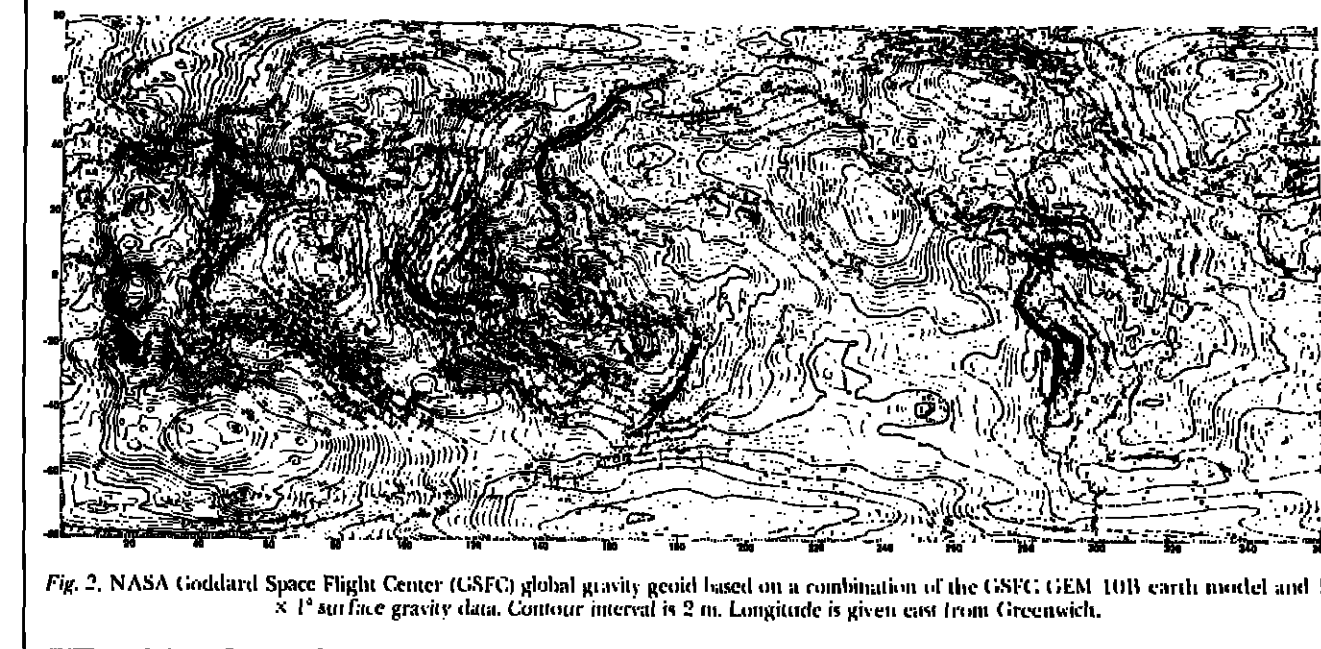


Fig. 2. NASA Goddard Space Flight Center (GSFC) global gravity geoid based on a combination of the GSCG (GEM 10B earth model and $1^\circ \times 1^\circ$ surface gravity data. Contour interval is 2 m. Longitude is given east from Greenwich.

1 and 2) is that the magnetic field has been recorded directly with flux gate, proton precession, and alkali vapor magnetometers, while the gravity field has been derived indirectly from resolving discrepancies in the satellite's motion. These perturbations in motion were revealed by precisely recording elements of the satellite's position and velocity via optical, radar/laser ranging, or Doppler techniques. With the current field models (Langel et al., 1981), we have a 500 km horizontal resolution; however, the GRM should provide a 100 km wavelength resolution.

Earth orbiting satellites are advantageous for geopotential field determinations owing to their rapid, global coverage. With magnetic

measurements, speed is particularly important since the field changes in time as well as space. These earlier missions have had two unsatisfactory characteristics: They were relatively high (>400 km) and had elliptical orbits. The Geopotential Research Mission will record both the gravity and magnetic fields from a 100 km altitude circular orbit. This project will unify the measurement of these major geopotential fields by recording them simultaneously. We will discuss this planned project and its scientific rationale and present some simulations of the anticipated data set.

To ensure complete global coverage, GRM will consist of two satellites in a $90^\circ \times 0.1^\circ$ polar orbit of 100 km altitude for 6 months and be separated by up to 600 km (Figure 1). The operation altitude of 100 km was chosen to reconcile the scientific requirements for resolution with the engineering designs for the size of the fuel tanks. These parameters are firm mission requirements and these spacecraft have been designed for fuel and drag considerations, to fully meet these parameters. Gravity values will be derived from recording the range-rate between the two spacecraft, while the lead vehicle will project a boom containing scalar and vector magnetometers and the attitude-determining sun sensors and star cameras.

The time of launch of this proposed new mission, if approved, will most likely be 1989-1990.

Scientific Objectives

Because GRM will be launched into a 100 km circular polar orbit, it will measure both the gravity and magnetic fields of the earth with an accuracy and resolution not as yet obtainable from previous spacecraft. Unlike ear-

lier satellite altimetry missions, GRM will record gravity field data over continental as well as oceanic areas. Having this accurate and detailed data set will allow us to study many regional scientific problems of the solid earth.

Both gravity and magnetic reference field models (such as in Figures 2 and 3) will be greatly improved by the data from GRM. A more detailed gravity field model will allow additional refinements to be made for the determination of other satellite orbits (for example, of altimetry bearing satellites). Previously completed altimetry missions (e.g., Seasat and GEOS 3) could have their orbital parameters re-determined yielding a much more accurate estimation of sea-surface height. Together with the early improved geoid, this re-determination will render the sea-surface topography during those previous missions in unprecedented detail.

Further improvements to the global geoid datum will also be achieved. Since all geophysical anomaly fields are produced by removing the theoretical or reference field from the observed data, any significant improvement will result in refined anomaly mapping, particularly in the area of crustal studies.

Enhancement of the intermediate wavelength features of the gravity field models (100 km $\leq \lambda \leq 1000$ km) are important for our understanding of the mass variation in the upper mantle. These mass anomalies are probably related to the forces driving the lithospheric plates as well as indicating the physical

Article (cont. on p. 610)

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Patrick T. Taylor received a B.S. in geology (Michigan State) and an M.S. and Ph.D. in geophysics (Penn State and Stanford). At the Geophysics Branch at GSFC he is currently involved with the geologic interpretation of satellite geophysical data.

T. Keating graduated from Georgia Tech (B.E.E., 1956) and Catholic University (M.S.E.E., 1964). He has had additional graduate work in modern communications at George Washington University (1978 to 1981). He is the study manager for the GRM and works in the Advanced Mission Analysis Office of GSFC.

Werner D. Kahn holds a B.S. (1952) and M.S. (1953) in mathematics from the University of Illinois. Since 1960 he has worked at GSFC, specializing in geodynamics and ocean dynamics.

TABLE 1. Orbital Characteristics of Significant Satellites Used for Modeling the Earth's Gravity Field

Satellite	Inclination, degrees	Altitude, Range, km	Launch Date	Instrument	Approximate Accuracy
Agnes Rocket Body	69.01	912-926	1964	O	3-4"
Anna-1B	50.1	1076-1182	Oct. 1962	O	3-4"
BE-8	79.69	876-1075	1964	L, RR, O	5 cm/s 2 m, 5 cm/s 3-4"
BE-C	41.19	938-1322	1965	O	3-4"
Counter-1B	28.51	971-1211	1960	O	5 cm/s
DI-C	30.97	572-1553	1967	O	5 cm/s
DI-D	80.46	598-1890	1967	L, O	3-4"
Edo Rocket Body	47.21	1494-1682	1960	O	3-4"
GEOS-1	59.4	1116-2277	Nov. 6, 1965	L, RR, O	1-2 m, 5 cm/s 2"
GEOS-2	105.8	671-976	Jan. 11, 1968	L, RR, O	1-2 m, 1 m 5 cm/s 1-2"
GEOS-3	115.0	821-854	April 9, 1975	L	25 cm 3-4"
GOS	49.76	428-1294	1963	O	5 cm/sec
INJUN	66.82	880-996	1961	O	3-4"
Landsat-1	98.8	808-911	July 23, 1972	RR	5 cm/sec
Lapras	109.8	5888-5945	May 4, 1976	L	3-4"
Midas-1	95.83	3505-3729	1961	O	3-4"
OGO-2	87	413-1610	Oct. 1965	O	3-4"
OSCAR-7	89.7	867-1199	1966	O	3-4"
OV-13	144.27	1786-2092	1965	O	3-4"
SECOR-5	69.2	1135-2418	Aug. 10, 1965	O	3-4"
Sarosite	49.8	806-1108	Feb. 6, 1973	L	5 cm
PEOLE-1	15.0	395-465	Dec. 12, 1970	L, M	1 m 20"

M, Minitrack; L, Laser Ranging; R, Radar Range; RR, Doppler; and O, Optical.

TABLE 2. Satellites That Have Measured the Near-Earth Geomagnetic Field

Satellite	Inclination, degrees	Altitude, Range, km	Launch Dates	Instrument	Approximate Accuracy, nT
Sputnik 3	65	440-600	May-June 1958	Fluxgate	100
Vanguard 3	35	510-5750	Sept.-Dec. 1959	Proton	100
1963-38C	Polar	1100	Sept. 1963-Jan. 1974	Fluxgate (1 axis)	30-85
Comet 26	49	270-403	March 1964	Proton	Unknown
Comet 40	50	261-488	Oct.-Nov. 1964	Proton	22
1964-33C	90	1040-1089	Dec. 1964-June 1965	Rubidium	22
OGO-2	87	413-1510	Oct. 1965-Sept. 1967	Rubidium	6
OGO-4	86	412-908	July 1967-Jan. 1969	Rubidium	6
OGO-6	82	397-1098	June 1969-July 1971	Rubidium	6
Comet 921	72	270-403	Jan.-March 1970	Cesium	Unknown
Azur	103	384-5145	Nov. 1969-June 1970	Fluxgate (2 axis)	Unknown
Triad Magnet	Polar	750-832	Sept. 1972-present	Fluxgate	6
	Polar	352-561	Oct. 1979-June 1980	Fluxgate Cesium	2

From Langel (1980).

Article (cont. from p. 609)

cal characteristics of material in the asthenosphere. Shorter wavelength features ($\lambda < 100$ km) are indicative of the relative density, nature, and physical state of the lithosphere and crust. *Marsh and Marsh* (1976) and *Lambert* (1976) have correlated patterns in the global free-air gravity anomaly field with geological features, convection, and density inhomogeneities.

Magnetic field data obtained from GRM will enable comparable insights into our understanding of the solid earth. Both the vector and scalar magnetic field measurements will be used to further refine the geomagnetic reference field models, which are being continually improved (Figure 3). When the dense and precise GRM values are used alone or combined with the earlier reference field (*Langel et al.*, 1980) we will secure the time-varying magnetic field as never before. The time terms derived from this comparison can be applied to the commonly used International Geomagnetic Reference Field (*Peddie*, 1981) for the practical applications of magnetic survey reductions and compilation of aerally contiguous magnetic charts. In maintaining global reference field models it is necessary to record the magnetic field periodically in order to detect the unsteady temporal field variations.

Beyond the mechanics of refining the reference fields for chart reduction and navigational purposes, a great deal can be learned about the earth's interior from these studies. The energy distribution in the lower harmonics in the magnetic field are important parameters in studies of the earth's core (*Benton et al.*, 1979).

Shorter wavelength features ($\lambda \leq 500$ km) of the magnetic field, the so-called anomaly field, have been interpreted to be the result of: (1) intracrustal contrasts in magnetization (*Regan and Marsh*, 1982; *Mayhew et al.*, 1982); (2) variations in the thickness of the crustal magnetized layer by Curie isotherm depth (*Mayhew*, 1982) or crustal thickness (*Schuetzler and Allenby*, 1983); and (3) surface geological features (*Fry*, 1982).

With the greater increase in resolution of the magnetic anomaly field from the GRM it may be possible to derive a geological/tectonic framework for economic evaluation as well as solve significant plate-tectonic problems.

The gravity and magnetic field data sets obtained by GRM will be complementary and supplementary. Both will furnish information about the lithosphere, its bulk composition, physical properties, and state. But gravity data will be primarily sensitive to mantle inhomogeneities and processes while magnetic measurements will mainly reflect the nature of the core. These geophysical potential field results will truly enable us to study the entire earth. These data sets will be useful to the studies of the smaller scale ($\lambda \approx 100$ km) global-crustal features (such as foldbelts, rifts, structural basins, and oceanic fracture zones and seamounts).

Mission Profile

Two satellites (Figure 1) designated A1 and A2 will be launched by the shuttle into a polar ($90^\circ \pm 0.1^\circ$) orbit. The shuttle will launch the satellites at 275 km altitude, and they will self-deorbit to a lower operating altitude of 160 km to begin a 6-month scientific mission. The spacecraft will orbit the earth with a period of about 88.5 minutes, providing approximately 16.5 orbits per day. Precision tracking of both spacecraft will be determined by the Defense Mapping Agency satellite tracking network.

Magnetometers, star cameras, and sun sensors will be located on an aluminum boom projected about 4 m from the A1 spacecraft (Figure 1). This lead satellite will weigh about 2800 kg, which is 200 kg greater than A2.

Gravity values will be derived by measuring the Doppler or range rate between the two satellites. The Satellite-to-Satellite Doppler Tracking System, which consists of ultrasonic oscillators, multipliers, and horn antennas (the latter being mounted above and below each spacecraft (Figure 1)) will be able to detect small changes in the distance by measuring the Doppler frequency shift of the continuous wave of the 91 and 42 GHz signals.

These continuous wave signals are transmitted between the A1 and the A2 spacecraft. The Doppler-shifted signals will be compared to on-board reference frequencies, and consequently the variation in the range between the spacecraft will be determined. The range variation is continuously time-sampled. The time-sampled range variation provides a measure of the relative velocity between the spacecraft to a sensitivity of 10^{-6} m s $^{-1}$.

Separating gravity field perturbations on the satellite orbits from external influences (such as atmospheric drag, solar radiation pressure, and solar and lunar gravity) is accomplished by the Disturbance Compensation System (RCS, 1983). This system consists of a 14 cm diameter ball (called the "proof mass") within a 10 cm diameter spherical cavity. The spherical cavity is also an electrical capacitor. The position of the proof mass in the cavity is determined by the capacity relative to each axis. When the proof mass is off center, the resultant imbalance on the capacity is resolved into vector thrust commands to the appropriate thrusters of the propulsion system. The resultant movement of the spacecraft recovers the proof mass within the cavity. Since the spacecraft shields the proof mass from all perturbing surface forces, the mass responds only to variations in the gravity field. The two spacecraft positioned by the proof mass's response to gravity are themselves the instruments for detecting the gravity field. Interaction between the spacecraft and the earth's electromagnetic field could result in small, torque-like forces. These relatively weak electromagnetic field forces are countered by the reaction-wheel torque, thus eliminating any secular effects.

Distance between the spacecraft may be varied up to 600 km in order to more accurately measure specific harmonics of the gravity-field spectrum.

According to the mission plan, the maximum distance between profile crossings of the equator will be about 7 km. The design goals for the earth's gravity field analysis are a 1 milligal accuracy in field measurement with a 5 cm geoid height resolution; both will be resolved to less than a degree (100 km).

Magnetic field determination should have the same spatial resolution with scalar field accuracy of 2 nT and vector components to 20 nT. All stated accuracies are at the 1-sigma level.

Data Simulation

In an effort to assess the improvements to be realized in the gravity and magnetic fields from the GRM, simulations were conducted for recovery of both fields using the mission nominal orbit altitude and inclination. These studies provide an estimate of the geological and geophysical interpretations which can be derived from the GRM's data (*see Douglas et al.*, 1980; *Jekeli and Rapp*, 1980; *Picane and Vinnova*, 1980; *Columbo*, 1981).

Magnetic anomaly field simulations were aided by the availability of a digitized version of the new Composite Magnetic Anomaly Map of the United States (*Zietz et al.*, 1982), obtained from the Phoenix Corporation by Goddard Space Flight Center under a proprietary contract. *Schuetzler et al.* (1983) determined that there was a long wavelength component (~ 5000 km) in this U.S. magnetic anomaly map. Therefore before extrapolating these data for higher altitudes ("upward-continuing" the data), a long-wavelength trend surface was removed by applying a general, two-dimensional polynomial surface (with ninth-order terms in longitude and fourth-order terms in latitude with cross terms; this is equivalent to 45 degrees of freedom). Theoretically, we should be able to apply the long-wavelength (low pass) filter to the data set either before or after the upward continuation. In practice, however, if this filter were applied to the data after continuation, the long-wavelength components would be completely dominated by the field and the shorter wavelength features would be harder to detect. The magnetic anomaly data points were selected at 0.2° intervals and upward-continued to 160 km altitude by a mathematical algorithm of *Huhtacharya* (1977), modified by J. Phillips of the U.S. Geological Survey. At this altitude the field points were plotted at a

0.2° grid interval and contoured (Figure 4). Profiles were taken through this data and one is presented in Figure 5.

When we compare the upward-continued field (Figure 4) with a recent U.S. MagSAT anomaly map at 320 km altitude (Figure 6) (*Mayhew and Galtier*, 1982), we can, not surprisingly, note the greater resolution of the lower orbiting satellite. Theoretical modeling studies, however, indicate that an even greater resolution of the magnetic anomaly field should be apparent. Differences in resolving power between the idealized models and the upward-continued U.S. magnetic anomaly field may be a result of a less than optimum field being removed from the latter. Further efforts are under way to define a more appropriate regional field for the digitized data base for the United States.

For gravity-field simulations, a free-air anomaly map was computed at 160 km altitude from the U.S. Geological Survey data set (Figure 7) [U.S. Geological Survey, 1982] using the same upward-continuation algorithm used with the magnetic field simulation. More efficient computer usage required that the original 4 km spacing between data points be increased to 20 km. Since we were computing the field at an altitude of 8 times the data interval, this was not considered to be a serious problem. These gravity values were plotted at the GRM nominal altitude (160 km) and at the nominal MagSAT height (320 km). This comparison was used to suggest the increasing anomaly resolution to be expected from the lower-orbiting GRM.

In an effort to assess the geological significance of the potential field data anticipated from this proposed mission, a simplified tectonic map was superimposed on the magnetic field simulated at 160 km altitude (Figure 8). Since the magnetic measurements were not reduced to the pole, a strict spatial correlation cannot be made.

Conclusions

The proposed GRM would accomplish the task of the satellites listed in Tables 1 and 2. Gravity and magnetic field modeling together with crustal anomaly studies would be the two major geophysical disciplines to benefit from this mission. GRM data would also be relevant to more specialized topics such as earth-core modeling, geoid ocean surface separation, and ocean current investigations. Engineering studies and data simulations

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Cover. Simulation of the launching of a Geopotential Research Mission (GRM) satellite from the space shuttle. The GRM mission has been proposed by the National Aeronautics and Space Administration (NASA) for launching by the end of the decade. (See article, this issue.) Photo courtesy of Patrick T. Taylor, NASA Goddard Space Flight Center, Greenbelt, MD 20771.)

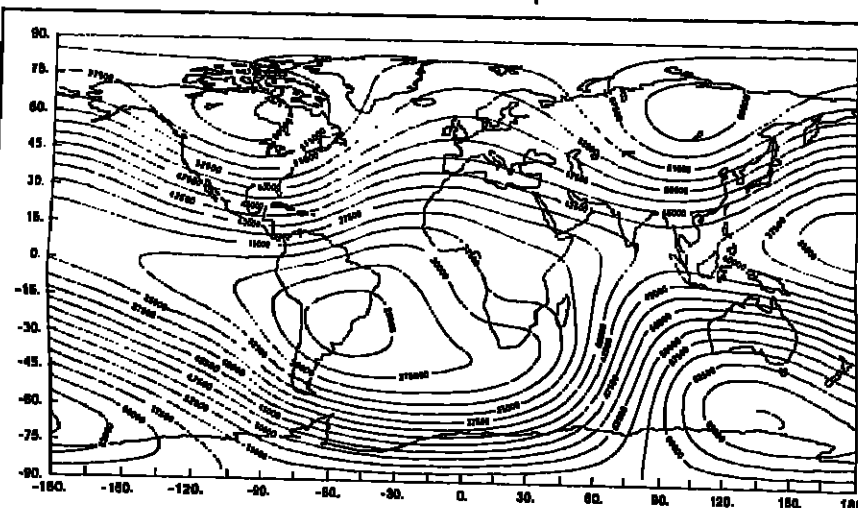


Fig. 3. MagSAT-derived total-magnetic field map computed at the earth's surface. Contour interval is 2500 nT.

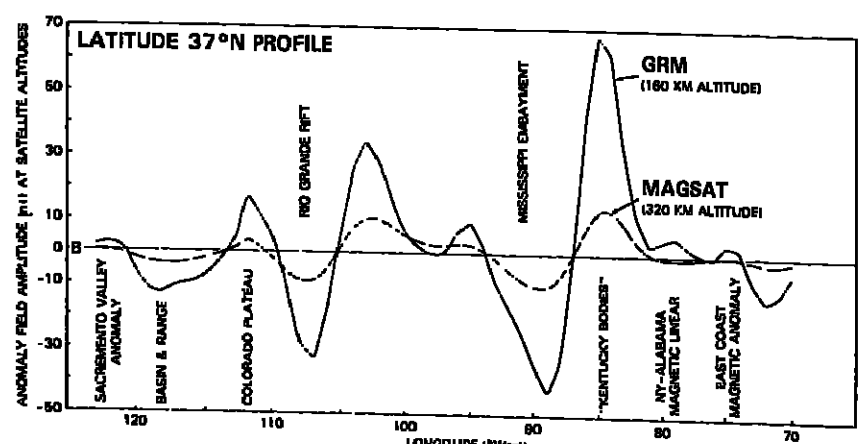


Fig. 5. Magnetic anomaly field profiles along 37° north latitude. Data from Figures 4 and 6. Prominent tectonic/geophysical features are given for reference.

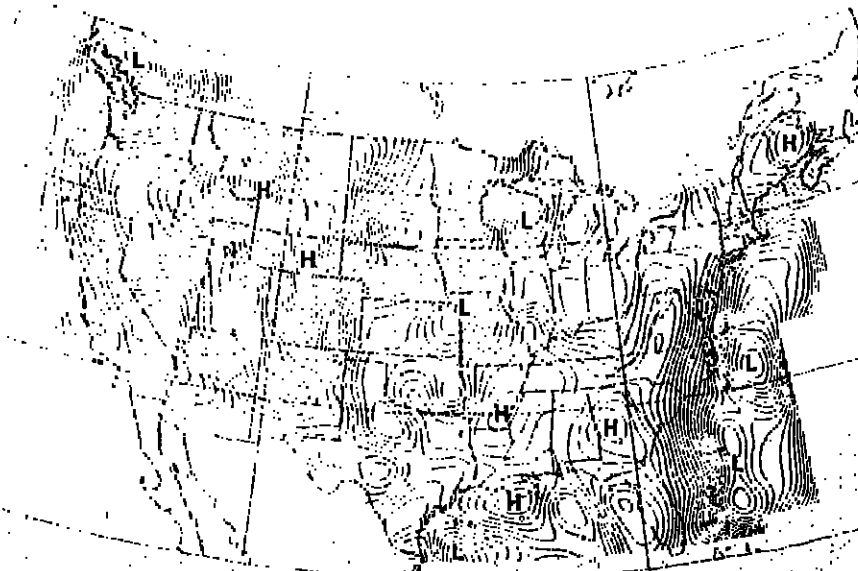


Fig. 7. Free-air gravity field data over the United States (data from U.S. Geological Survey) computed at 160 km altitude. Contour interval is 1 mgal.

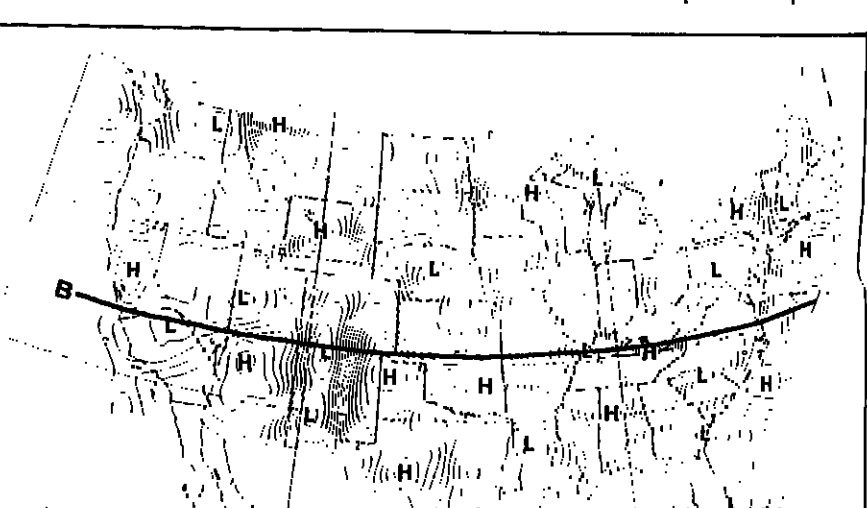


Fig. 4. U.S. Composite Magnetic Anomaly Map continued upward to 160 km altitude. Contour interval is 5 nT.

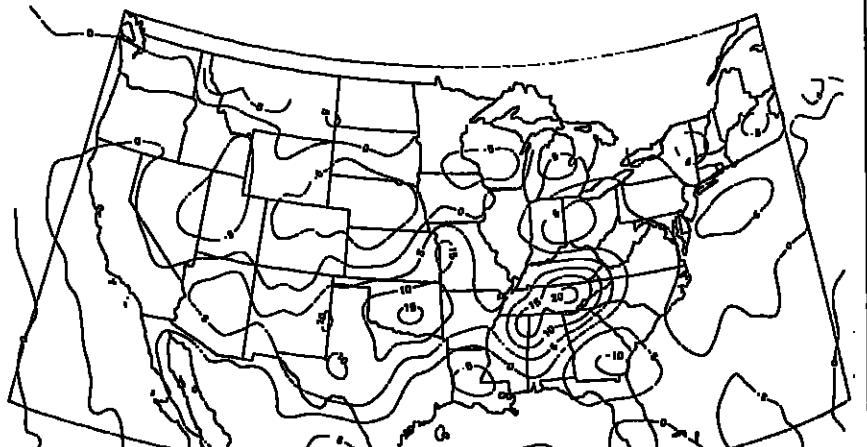


Fig. 6. Magnetic anomaly map of the United States computed at 320 km altitude (from Mayhew and Galtier, 1982). Contour interval is 5 nT.

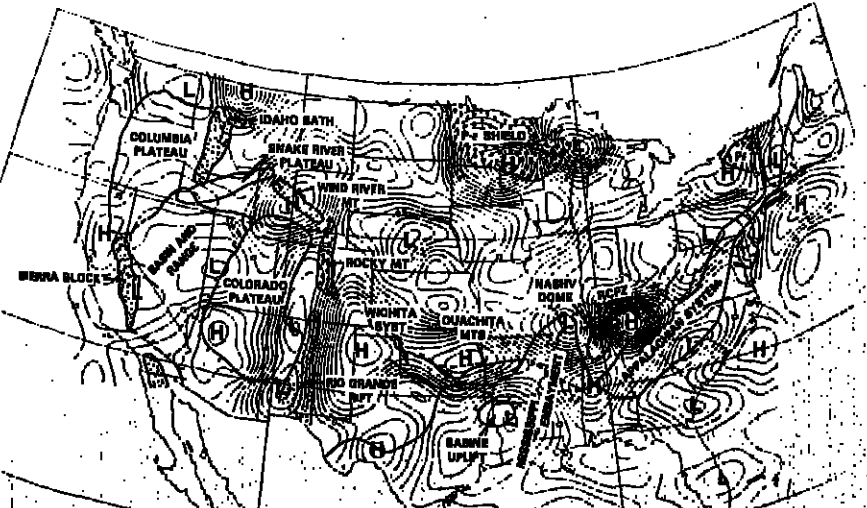


Fig. 8. Magnetic anomaly field data over the United States (Figure 4) with major geological tectonic features superimposed (based on King [1977]).

have revealed that the mission goals (gravity field to 2 milligals, geoid to 10 cm, and magnetic field anomaly map to 1 nT with a 20 km resolution) are realistic and should be achievable.

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News

Marine Sulfur Sources

The world's oceans have been known for several years to be a significant source of sulfur in the atmosphere. Current estimates place sulfur from marine sources at about 50% of the atmospheric sulfur attributable to human sources. However, recent work published in *Geophysical Research Letters* (GRL) indicates that estimates need to be refined to account for seasonal variations in the production of marine sulfur and poor understanding of air-sea inter-lake processes.

Dimethyl sulfide (DMS) represents more than 90% of the volatile sulfur compounds found in the ocean. Joel D. Cline and Timothy S. Bates, in the October issue of GRL, show that production of DMS in the central equatorial Pacific is strongly correlated with chlorophyll "a," a measure of phytoplankton abundance, which is in turn associated with equatorial upwelling. The concentrations of DMS (Cline and Bates found from 0°N to 0°S and from 148°W to 170°E about 2% of the world's ocean surface area) would yield about 1% of the entire annual flux of sulfur from the ocean to the atmosphere.

The findings of Cline and Bates in the equatorial Pacific accord with those of M. O. Andreae and H. Raundornck (*Science*, August 10, 1983), who sampled DMS there and elsewhere in the Pacific and Atlantic oceans. However, since DMS in the oceans is correlated with biological production and possibly phytoplankton speciation, its contribution to atmospheric sulfur will likely vary with changes in the seasons and geographical areas.

"Only a small fraction of the world oceans have been investigated and little progress has been made on the seasonal variations at high latitudes," says Cline and Bates. In addition, "Oceanic fluxes . . . are currently based on a large number of assumptions, chief among them . . . the validity of the stagnant film boundary layer model as it applies to a reactive species such as DMS."

The work of Cline and Bates was supported by the federal government's interagency National Acid Precipitation Assessment Program, part of whose program is measuring the relative contributions of natural and anthropogenic sources of atmospheric sulfur.

This summary of a paper in the current Geophysical Research Letters was contributed by Joel D. Cline.

Venusian Lightning

Laboratory simulations by William J. Borucki and coworkers (*Geophysical Research Letters*, October 1983) indicate that the inability of satellite instruments to detect optical pulses of Venusian lightning is due to the weakness of the signal rather than its spectral characteristics.

Evidence of lightning activity in the Venusian atmosphere has been obtained by instrumentation aboard the Pioneer Venus Orbiter (PVO) and Venera 9, 11, and 12 spacecraft. However, a search by Borucki and coworkers for optical pulses radiated by Venusian lightning was unsuccessful. The search used the star sensor aboard the PVO to cover the geographical region from 70°N to 50°S latitude and from 240° to 60° longitude (IAU coordinates). B. Vonnegut and R. E. Orville have suggested that Venusian lightning might emit very little radiation of wavelengths greater than 480 nm. This would reduce the possibility of detecting and mapping lightning activity using many solid-state sensors because these sensors are most sensitive in the red rather than in the blue spectral region.

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ity of detecting and mapping lightning activity using many solid-state sensors because these sensors are most sensitive in the red rather than in the blue spectral region.

To test the hypothesis of spectral limitation, the lightning facility at the NASA Langley Research Center was used to produce an electrical discharge across a 5-mm gap in a special gas mixture to simulate lightning in the Venusian atmosphere. Both aluminum and tantalum electrodes were used to assess the contribution of electrode radiation.

A comparison of the air spectrum from the capacitor discharge with that obtained by Orville and L. E. Salamea for terrestrial lightning showed that all the prominent lines in the terrestrial spectrum are also found in the spectrum from the capacitor discharge. Because the same lines appear with approximately the same relative intensity in both spectra, the capacitor discharge produces a useful simulation of terrestrial lightning, nevertheless such a simulation must be regarded as purely exploratory.

A comparison of the spectra obtained from the simulation with those obtained by Orville and Salamea for terrestrial lightning showed that (1) nitrogen lines are dominant in the terrestrial spectrum, but are not found in the simulated Venusian spectrum; (2) carbon lines are prominent in the simulated Venusian spectrum, but do not appear in the terrestrial spectrum; (3) oxygen lines and the hydrogen-H α line are present in both spectra.

A comparison of simulated and predicted spectra shows that most of the lines predicted are present in the experimental spectrum. A comparison of the predicted radiant fluxes from line radiation in the blue region (i.e., 350-400 nm) and the red region (600-900 nm) shows that the flux in the red region is approximately 5 times larger than that in the blue region.

It is fortunate that Venusian lightning is expected to radiate strongly in the region of 600 to 900 nm because the in situ measurements by V. I. Moroz et al. of atmospheric transmission and the radiative transfer calculations by M. A. Williams et al. show that very little radiation at wavelengths shorter than 500 nm will penetrate the Venusian atmosphere and clouds. In particular, lightning observed through the Venusian atmosphere will have a reddish hue. Consequently, the reason the star sensor did not detect the Venusian lightning was not its spectral characteristics; rather, it seems likely that the low flashing rate of the low intensity produced a signal level that was too low to be detected.

This summary of a paper in the current Geophysical Research Letters was contributed by William J. Borucki.

Planetary Instruments

November 30 is the deadline for submitting proposals for the National Aeronautics and Space Administration's (NASA) planetary instrument definition and development program (PIDDP). Proposals arriving after the deadline will be reviewed next year. Formerly called the planetary instrument definition program, PIDDP aims "to define and develop new instruments and instrument components or improved versions of existing instruments so that they may be proposed in response to future announcements of flight opportunities as candidate experiments that would not

need further development if selected."

NASA encourages new proposals that focus on support for studies aligned with future mission plans of the Solar System Exploration Division (SSE) (May 24, 1983, p. 386), and February 15, 1983, p. 65). Among the missions recommended by the Solar System Exploration Committee (SSEC) are the Venus Radar Mapper, the Mars Science/Climate Orbiter, the Comet Rendezvous/Asteroid Flyby, and Titan Probe/Radar Mapper. Subsequent missions recommended include the Mars Aeronomy Orbiter, the Venus Atmosphere Probe, the Mars Surface Probe, and the Lunar Geoscience Orbiter.

Those studies that are continuations of work already reviewed, accepted, and funded will continue to receive support in fiscal year 1984 at levels previously negotiated, unless the investigator requests a change in scope or an increase in the funding level. However, investigations of all continuing studies should note that fiscal 1984 will be the last year of approval for these ongoing efforts. Completely new proposals will be required for support in fiscal 1985 and beyond.

For additional information, contact Fred Veselich, Code EL-4 (PIDDP), NASA, Washington, DC 20546.

Geophysicists

Rear Admiral John D. Bossler, a member of the National Oceanic and Atmospheric Administration (NOAA) Corps and secretary of the AGU Geodesy section, is the new director of the National Oceanic and Atmospheric Administration's charting and geodetic services. Bossler was director of NOAA's National Geodetic Survey.

Frank Stehli, dean of geosciences at the University of Oklahoma, has been appointed chairman of the Continental Scientific Drilling Committee of the Board on Earth Sciences, National Research Council.

Recent Ph.D.'s

Eos periodically lists information on recently accepted doctoral dissertations in the disciplines of geophysics. Faculty members are invited to submit the following information on institution letterhead, above the signature of the faculty advisor or department chairman: the dissertation title, author's name, name of the degree-granting department and institution, and month and year degree was awarded. If possible include the current address and telephone number of the degree recipient (this information will not be published).

The Border Range Mafic and Ultramafic Complex: Petrologic Correlation of an Intracrustal Volcanic Island Arc, Laure E. Burns, Dept. of Geophysics, School of Earth Sciences, Stanford Univ., September 1983.

Crustal Shear Velocity Modeling in Nevada: A Study of Broadband Multi-Node Surface Waves, Kin Yip Chiu, Dept. of Geology and Geophysics, Univ. of California, Berkeley, December 1983.

Helium, Neon, Water, and Carbon in Volcanic Rocks and Gases, Robert Joseph Poreda, Graduate Dept. of Scripps Institution of Oceanography, Univ. of California, San Diego, December 1983.

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POSITIONS WANTED

Physical Chemist. Ph.D. specialized in isotopic geochemical kinetics. Research would consider challenging opportunity. P.O. Box 018 American Geophysical Union, 2000 Florida Avenue, N.W., Washington, D.C. 20009.

Geochemical/Geologist. Ph.D., 6 years of experience in trace element geochemistry, volcanology, hydrothermal processes, water-rock interaction and field techniques on land or at sea. Also teaching experience at university level. Seeks Research/Teaching/Industrial position in the Washington, D.C. area. Resume on request. American Geophysical Union, 2000 Florida Avenue, N.W., Box 018, Washington, D.C. 20009.

POSITIONS AVAILABLE

Geochemical/Alfred University. Invites applications for a tenure track position in geology commencing August 1984. We seek candidates with a Ph.D. who have a commitment to undergraduate education, are able to teach a variety of courses in geology which might include geochemistry, mineralogy and petrology, and have a strong interest in working closely with students. Alfred University has an enrollment of approximately 2100 students and is located in a rural setting in western New York state. There are approximately 23 geology majors and the department competes in a strong environmental studies program. Excellent academic and research facilities are available on campus in conjunction with the New York State College of Ceramics which, on campus of Alfred University, has a VAX 11/780 computer is available. Interviews will be conducted at the G.S.A. meeting in Indianapolis at Alfred University. Interested candidates should send a resume and three letters of recommendation to Dr. M.W. Webb, Chairman, Division of Physical Sciences, Alfred University, Box 832, Alfred, N.Y. 14802. Telephone (607) 871-2298.

Alfred University is an equal employment, affirmative action employer and encourages applications from women and other minorities.

North Carolina State University/Marine Chemist. The Department of Marine, Earth, and Atmospheric Sciences invites applications for a 9 month, tenure track position at the assistant or associate professor level. The candidate must have a Ph.D. and will be expected to interact with various research programs within the department such as radiocarbon, stable isotope and trace metal geochemistry, sedimentology, ocean circulation, air-sea interaction, and biological oceanography. Responsibilities include conducting a 9 month research program as well as teaching and advising graduate students. Applicants should forward a resume and the names of at least three references to: Dr. David J. DeMaster, Chairman, Search Committee, P.O. Box 5068, North Carolina State University, Raleigh, NC 27650. Application material should be sent by November 30, 1983.

North Carolina State University is an equal opportunity/affirmative action employer.

NASA NSDD/ACQUISITION SCIENTISTS. Sigma Data Services Corp., a M/A-COM Company, as contractor operating the National Science Data Center (NSDD) at NASA/GSFC, has immediate openings for scientists in the following disciplines:

—Astronomy/Solar Astronomy/Astrophysics (especially X-ray)

—Remote sensing/Meteorology/Atmospheric Sciences

—Linear/Planetary Geology/Petrology

—Magnetospheric Physics

M.S. required, Ph.D. preferred. Candidates should have experience in analysis of data from spacecraft experiments in their subject area. Working knowledge of FORTRAN is required. Incumbents will serve as acquisition agents for data archived at NSDDC, interface with investigators, and engage in data synthesis efforts, and the generation of data catalogs. Research opportunities available. Send resumes to Dr. H. K. Hills, Sigma Data Services Corp., NASA/GSFC Code 801, Greenbelt, Md. 20771 or call (301) 344-8105.

Deputy Department Head/Texas A&M University.

The Department of Oceanography in the College of Geosciences at Texas A&M University is seeking a deputy department head to assist in the academic and administrative functions in the Department. Duties will involve 75 percent administration and 25 percent research or teaching on a 12 month appointment basis. This is a tenure track position and will be filled at an academic level commensurate with the experience of the applicant. Applicant must have demonstrated administrative ability, an established record in research and teaching at both undergraduate and graduate levels of Oceanography. Closing date for applications is 1 November 1983. Effective date of this appointment will be 1 January 1984.

TAMU is an equal employment/affirmative action employer.

University of Alaska/Exploration Geophysicist—

Seismic Stratigrapher. Applications are invited for a tenure-track teaching/research position in the Geology/Geophysics Program of the College of Environmental Sciences. Prime responsibilities will be to teach graduate and undergraduate students in the use of state-of-the-art techniques in petroleum exploration geophysics. The successful applicant will also develop an innovative research program to complement the growing petroleum geology curriculum. Doctorate is required. Industry experience in hydrocarbon exploration and, in particular, the use of seismic reflection data to interpret stratigraphy and facies is desired. The three-month faculty position is open starting in January 1984. Rank and salary commensurate with qualifications and experience. Resumes and at least three references should be submitted to Dr. Juan G. Roederer, Director, Division of Geology, University of Alaska, Fairbanks, Alaska 99701. Applications will be accepted until December 15, 1983 or until position is filled.

Your application for employment with the University of Alaska may be included in Public Disclosure if you are selected as a finalist.

The University of Alaska is an EO/AA employer and educational institution.

University of Wisconsin-Parkside/Tenure-track Position. The Geology Program at the University of Wisconsin-Parkside invites applications for a tenure-track position at the assistant-professor level to begin in August 1984. The successful applicant will be expected to teach undergraduate courses in one or more of the following areas: hydrogeology, low temperature geochemistry, environmental geology; carry out a productive research program in either petrology and/or mineralogy; and share the teaching of introductory geology courses. The Ph.D. or equivalent is required. Submit a resume, transcripts, 3 letters of reference, and a statement of research and teaching interests by January 31, 1984 to:

Professor Gerald A. Fowler

Geology Program

University of Wisconsin-Parkside

Box No. 2000

Kenosha, Wisconsin 53141

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Applications should include a curriculum vitae, copies of one or two recent publications and the names of at least three referees. They should be sent by January 15, 1984 to Professor G.M. Phillips, Chairman, Search Committee, Department of Earth and Planetary Sciences, The Johns Hopkins University, Baltimore, Maryland 21218.

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Atmospheric Dynamics and Planetary Physics/The Johns Hopkins University. The Department of Earth and Planetary Sciences is seeking a tenure-track faculty appointment in each of these areas, one in July 1984 and the other in July 1985. The Atmospheric Dynamics position will be at the Assistant Professor level. The successful applicant will have a demonstrated capacity for innovative research with preferably, post-doctoral experience. The appointment in Planetary Physics will be made at a level commensurate with the applicant's qualifications. The successful candidate for the Atmospheric Dynamics position should have a high international reputation for research accomplishments. Women and minority candidates are especially encouraged to apply. It is expected that the appointee will develop programs in teaching and research that will complement the activities of the present groups in geophysical fluid dynamics and geophysics, whose research interests include turbulence, waves, air-sea interactions, stratified flow dynamics, convection, mesoscale meteorology, dynamics of the earth's interior, and volcanology. The appointee in planetary physics will be encouraged to interact closely with scientists at the Space Telescope Science Institute, which is on campus. The University is a member of the University Corporation for Atmospheric Research.

Applications should include a curriculum vitae, copies of one or two recent publications and the names of at least three referees. They should be sent by January 15, 1984 to Professor G.M. Phillips, Chairman, Search Committee, Department of Earth and Planetary Sciences, The Johns Hopkins University, Baltimore, Maryland 21218.

The Johns Hopkins University is an Equal Opportunity Employer.

NASA NSDD/ACQUISITION SCIENTISTS. Sigma Data Services Corp., a M/A-COM Company, as contractor operating the National Science Data Center (NSDD) at NASA/GSFC, has immediate openings for scientists in the following disciplines:

—Astronomy/Solar Astronomy/Astrophysics (especially X-ray)

—Remote sensing/Meteorology/Atmospheric Sciences

—Linear/Planetary Geology/Petrology

—Magnetospheric Physics

M.S. required, Ph.D. preferred. Candidates should have experience in analysis of data from spacecraft experiments in their subject area. Working knowledge of FORTRAN is required. Incumbents will serve as acquisition agents for data archived at NSDDC, interface with investigators, and engage in data synthesis efforts, and the generation of data catalogs. Research opportunities available. Send resumes to Dr. H. K. Hills, Sigma Data Services Corp., NASA/GSFC Code 801, Greenbelt, Md. 20771 or call (301) 344-8105.

Deputy Department Head/Texas A&M University.

The Department of Oceanography in the College of Geosciences at Texas A&M University is seeking a deputy department head to assist in the academic and administrative functions in the Department. Duties will involve 75 percent administration and 25 percent research or teaching on a 12 month appointment basis. This is a tenure track position and will be filled at an academic level commensurate with the experience of the applicant. Applicant must have demonstrated administrative ability, an established record in research and teaching at both undergraduate and graduate levels of Oceanography. Closing date for applications is 1 November 1983. Effective date of this appointment will be 1 January 1984.

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University of Alaska/Exploration Geophysicist—

Seismic Stratigrapher. Applications are invited for a tenure-track teaching/research position in the Geology/Geophysics Program of the College of Environmental Sciences. Prime responsibilities will be to teach graduate and undergraduate students in the use of state-of-the-art techniques in petroleum exploration geophysics. The successful applicant will also develop an innovative research program to complement the growing petroleum geology curriculum. Doctorate is required. Industry experience in hydrocarbon exploration and, in particular, the use of seismic reflection data to interpret stratigraphy and facies is desired. The three-month faculty position is open starting in January 1984. Rank and salary commensurate with qualifications and experience. Resumes and at least three references should be submitted to Dr. Juan G. Roederer, Director, Division of Geology, University of Alaska, Fairbanks, Alaska 99701. Applications will be accepted until December 15, 1983 or until position is filled.

Your application for employment with the University of Alaska may be included in Public Disclosure if you are selected as a finalist.

The University of Alaska is an EO/AA employer and educational institution.

University of Wisconsin-Parkside/Tenure-track Position. The Geology Program at the University of Wisconsin-Parkside invites applications for a tenure-track position at the assistant-professor level to begin in August 1984. The successful applicant will be expected to teach undergraduate courses in one or more of the following areas: hydrogeology, low temperature geochemistry, environmental geology; carry out a productive research program in either petrology and/or mineralogy; and share the teaching of introductory geology courses. The Ph.D. or equivalent is required. Submit a resume, transcripts, 3 letters of reference, and a statement of research and teaching interests by January 31, 1984 to:

Professor Gerald A. Fowler

Geology Program

University of Wisconsin-Parkside

Box No. 2000

Kenosha, Wisconsin 53141

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